



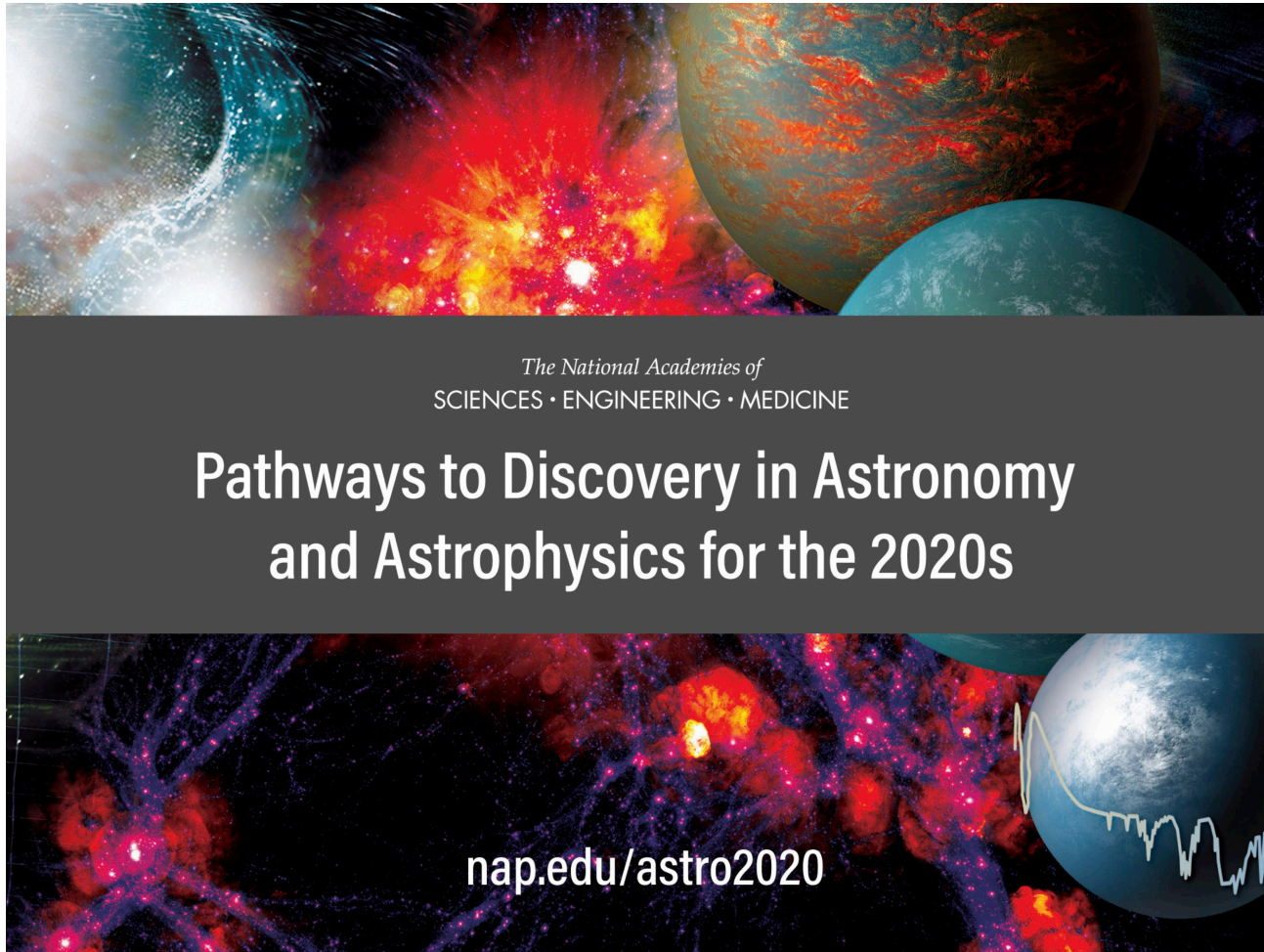
*The PRobe far-Infrared
Mission for Astrophysics*

PRIMA Overview



Jason Glenn, Principal Investigator, GSFC
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NASA Astrophysics Probes



Scope

Between MIDEX and Flagship: \$1.5B

Opportunity

Far-IR or X-ray Probe

Timeline

- Step 1 proposals: due 11/23
- Selection for Phase A: Last quarter 2024
- Concept Study reports: Due late 2025
- Selection for implementation: 2026
- Launch: 2032
- Community participation: *Ongoing!*

PRIMA Science



ORIGINS OF PLANETARY ATMOSPHERES



EVOLUTION OF GALACTIC ECOSYSTEMS

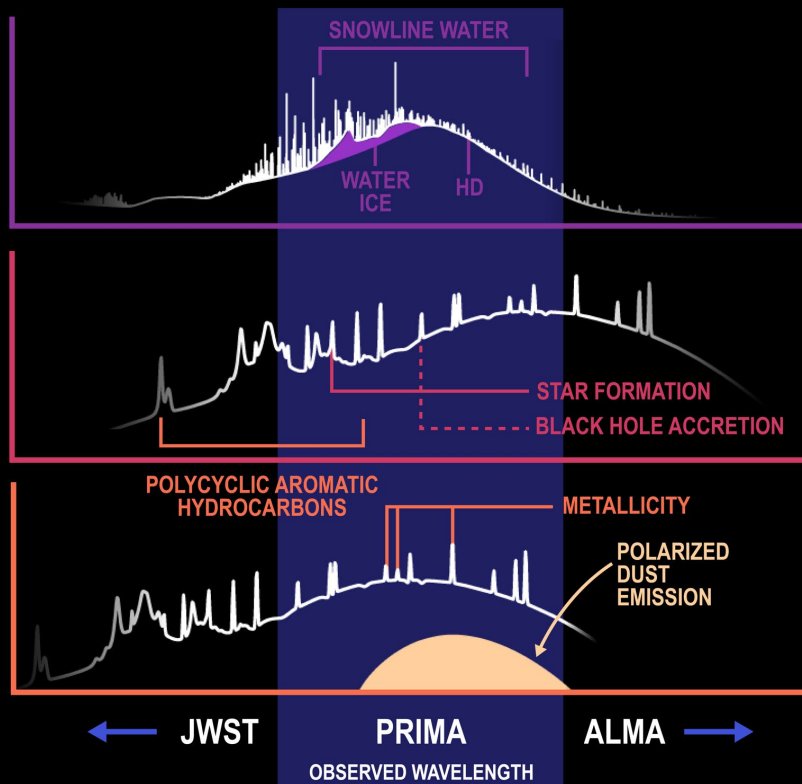


BUILDUP OF DUST AND METALS

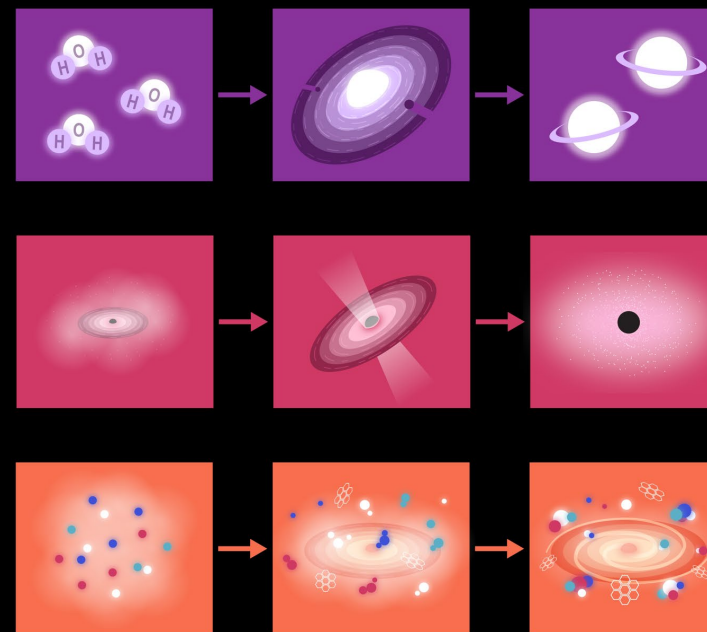


GENERAL OBSERVER AND GUEST INVESTIGATOR POTENTIAL

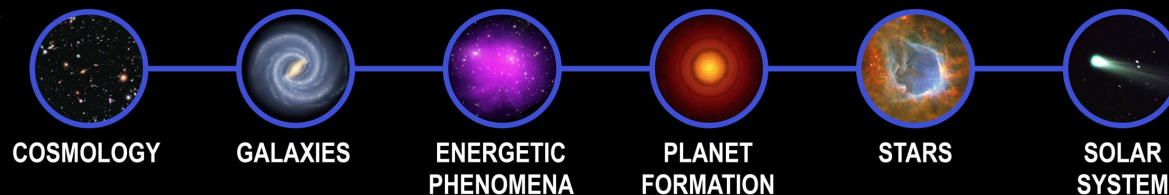
SPECTRAL MEASUREMENT



EVOLUTION



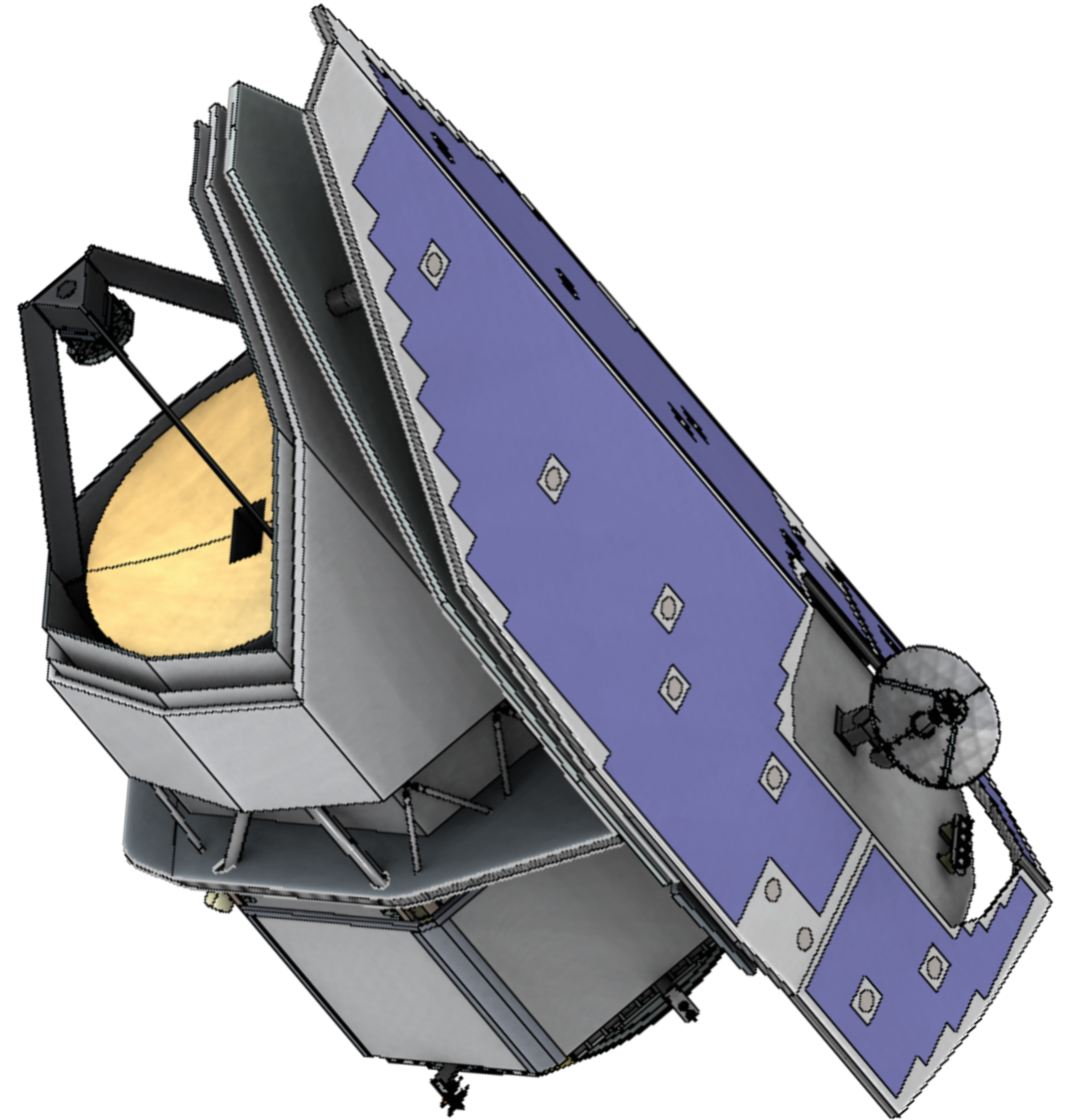
PRIMA uses the power of the far-infrared to see into the heart of dusty and obscured sources across cosmic time.



PRIMA's powerful and flexible capabilities enable general observer (GO) programs.

PRIMA Basic Facts

Observations	75% GO, 25% PI (→ GI)
Telescope	1.8 m, 4.5 Kelvin
PRIMAger (L. Ciesla)	Hyperspectral imaging 25-80 μm , R = 10 Imaging & polarimetry 91-261 μm , R = 4
FIRESS (M. Bradford)	Spectroscopy 24-235 μm , R > 85 High-Res mode R = 4,400 x ($\lambda/112\mu\text{m}$)
Detectors	Kinetic inductance detectors 11k total
Data	IPAC
Orbit	Earth-Sun L2
Launch	2032



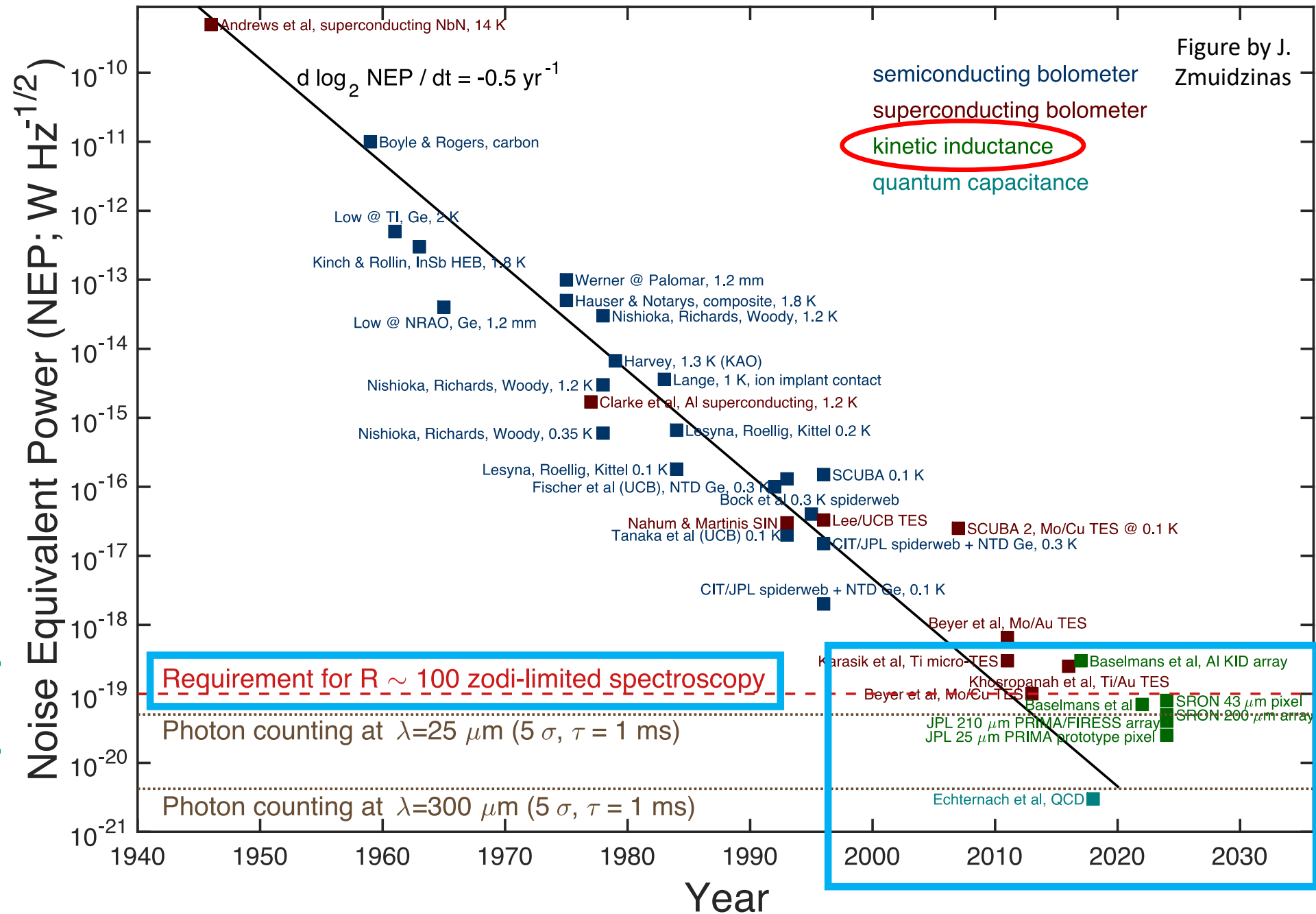


Why the Time is Now

Detector arrays have achieved unprecedented sensitivities and sizes

Sensitivities of far-IR detectors have doubled approx. every two years for 75 years!

Probe region of interest



Closing Thoughts

- Contact us if you have questions about sensitivity or observing calculations as you undertake your paper.
- If your science includes needs like dynamic range, include that in your papers so that we can consider them during Phase A.
- Community participation: *Ongoing!*
 - Talk to your colleagues about the electromagnetic spectrum on how far-infrared observations can enable their science.

Thank you to our JATIS guest editors: Naseem Rangwala and Matt Griffin!